

# Elliptic and Magneto-Elliptic instabilities

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NASA/JPL-Caltech



Marseille, September 2012

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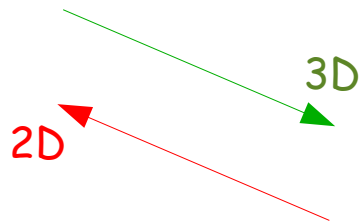
**Vortices are the fundamental unit of turbulent flow.**



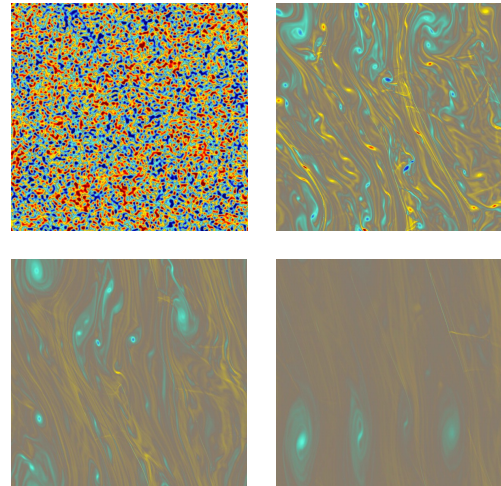
***"... the smallest eddies are almost numberless,  
and large things are rotated only by large eddies and not by small ones  
and small things are turned by small eddies and large"***

**da Vinci (1500), on *torbolenza***

# The energy cascade



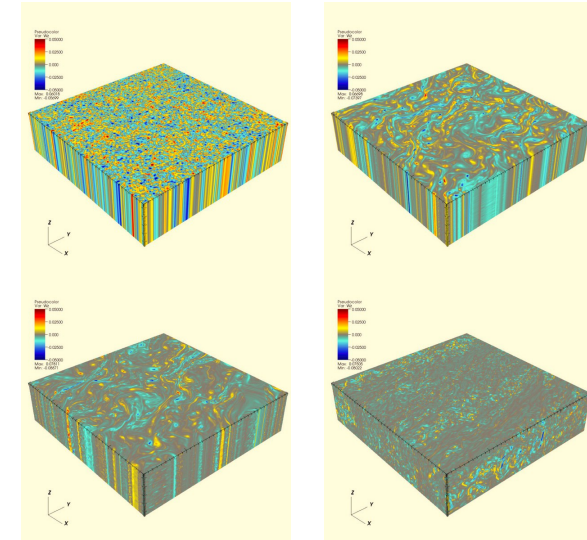
Shen et al. (2006).  
See also Batchelor (1967)



2D

**Inverse Cascade**

Eddies merge



3D

**Direct Cascade**

Eddies decay

Understanding the stability of vortices plays a fundamental role in understanding turbulence

# How stable are closed elliptic streamlines?

## Elliptic flow

$$U_x = -\chi \Omega_v y$$

$$U_y = \chi^{-1} \Omega_v x$$

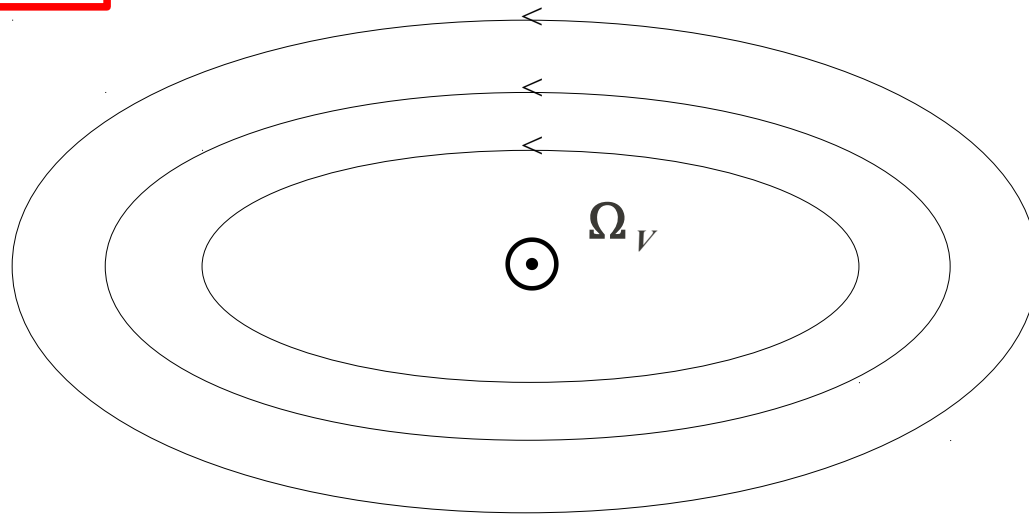
$\chi$  = aspect ratio

$\Omega_v$  = turnover frequency

Solve Euler equations for this flow

$$\partial_t u_i = -u_j \partial_j u_i - \rho^{-1} \partial_i p$$

$$\partial_i u_i = 0$$



# Inertial Waves

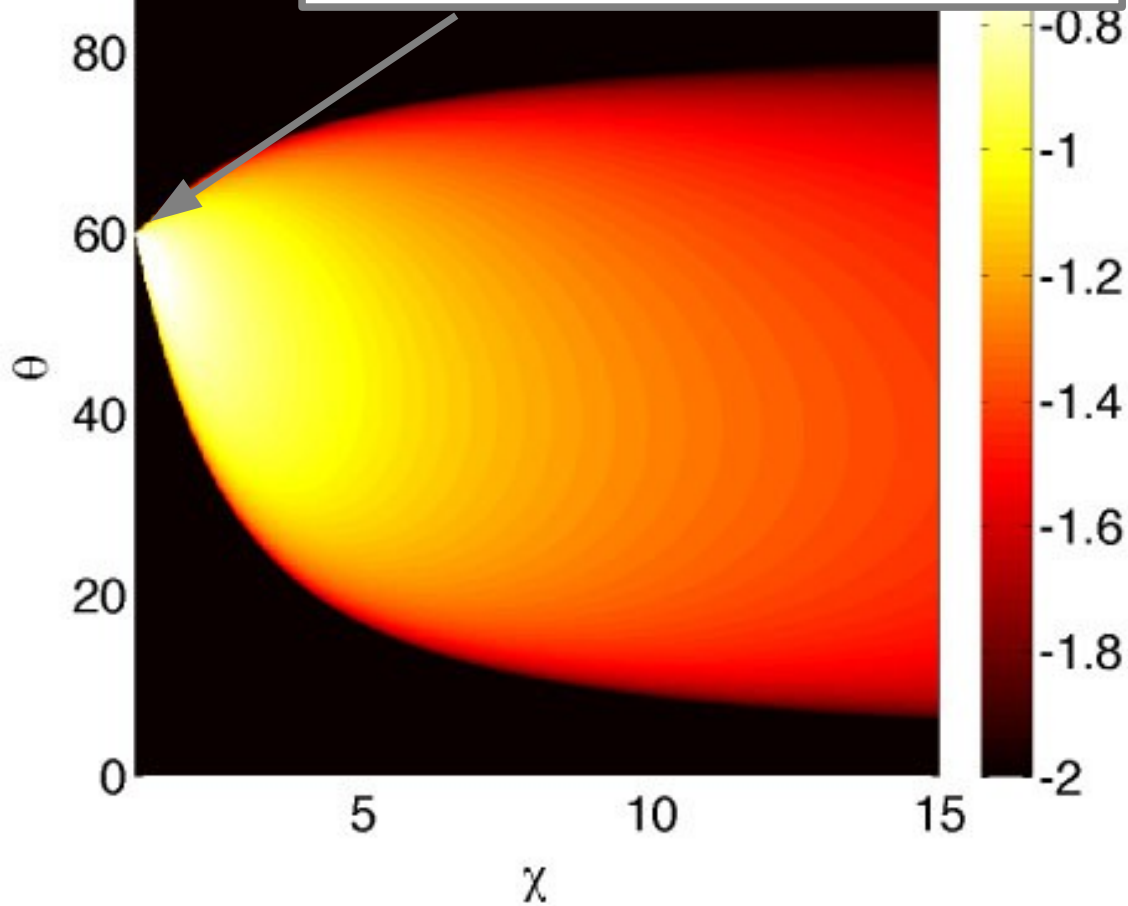


# Inertial Waves

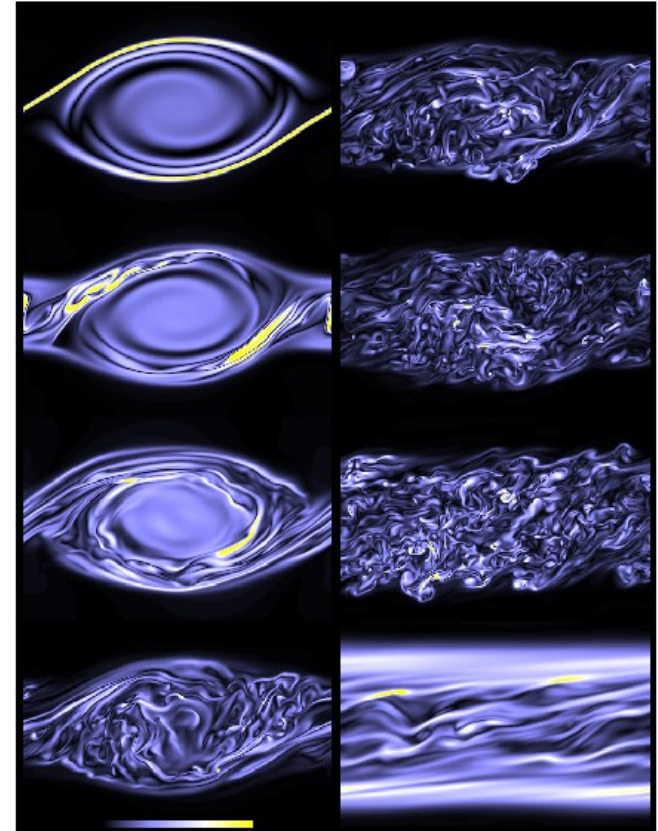


# Growth rates

No instability for circular streamlines.  
Thus, the instability is "elliptical".



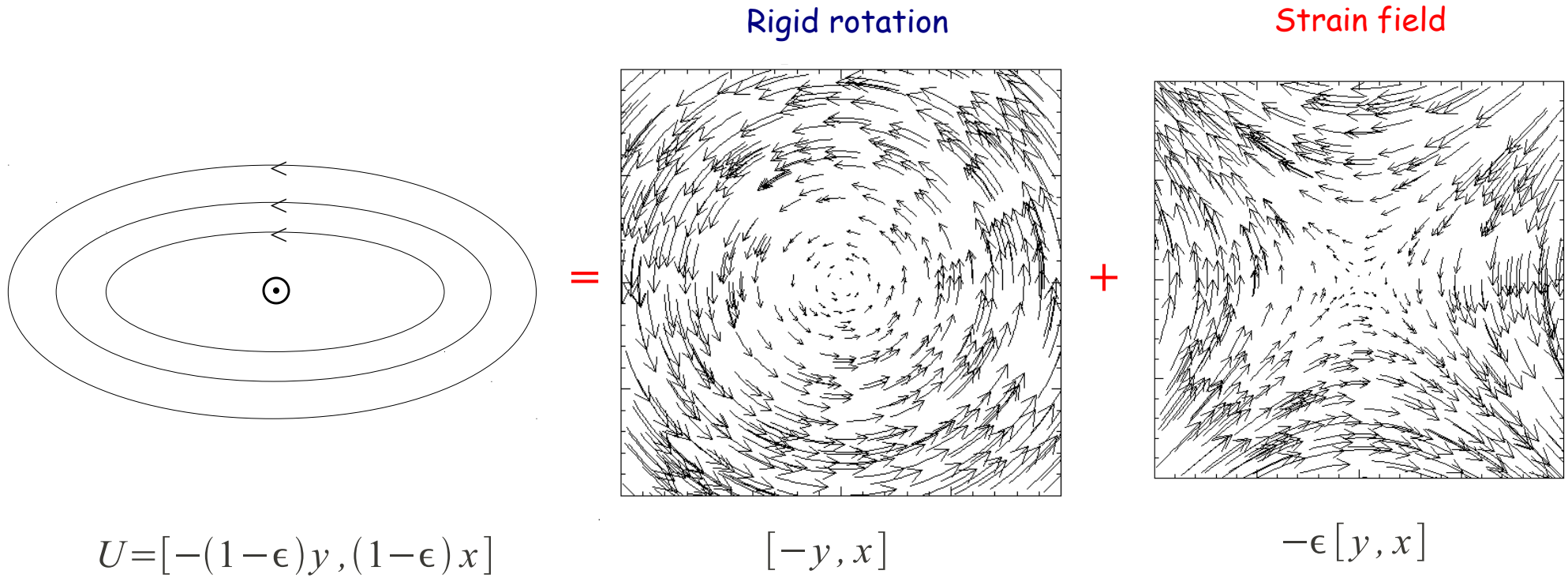
Lesur & Papaloizou (2009)  
After Bayly (1986)



Vortex coherence is destroyed  
Energy cascades forward and dissipates  
The flow relaminarizes

McWilliams (2010)

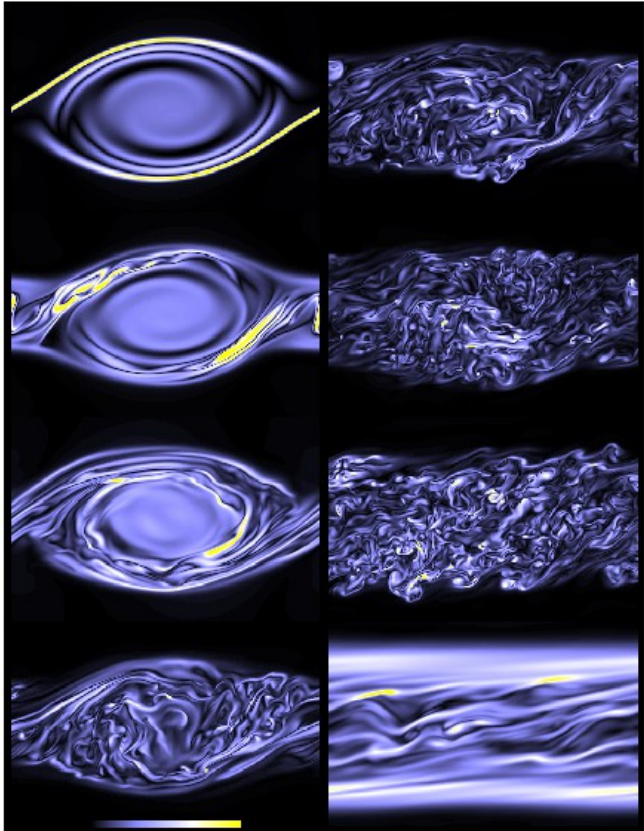
# Decomposing the motion



The helical oscillations  
are de-stabilized by the **strain field**.



## End Result



McWilliams (2010)

The instability is 3D

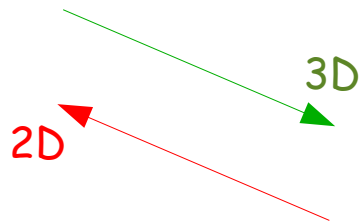
Generates 3D turbulence out of 2D motion

**secondary instability**

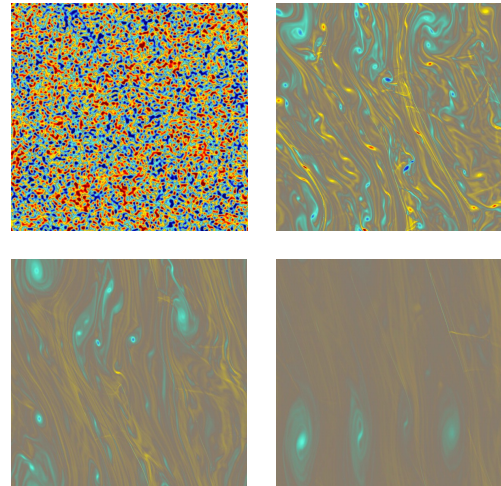
A stirring or primary instability (RT, KH)  
generates the first eddies.

The elliptic instability breaks them,  
leading to the direct cascade.

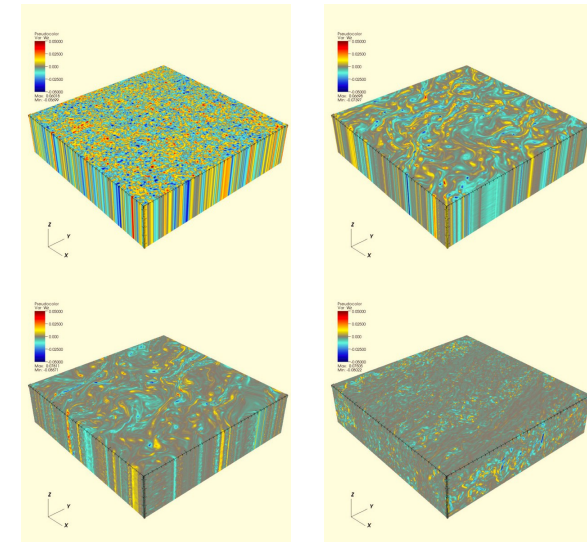
# The energy cascade



Shen et al. (2006).  
See also Batchelor (1967)



2D



3D

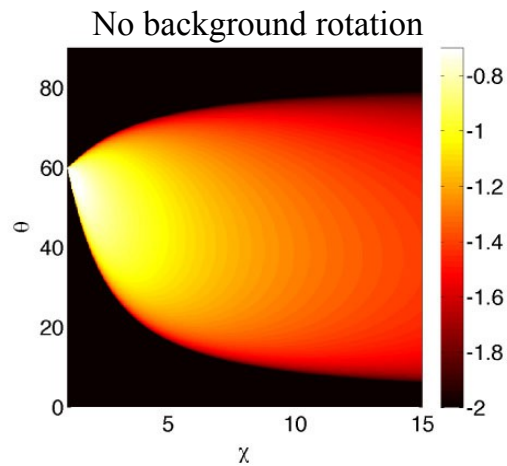
## Inverse Cascade

No elliptic instability  
Eddies merge viscously

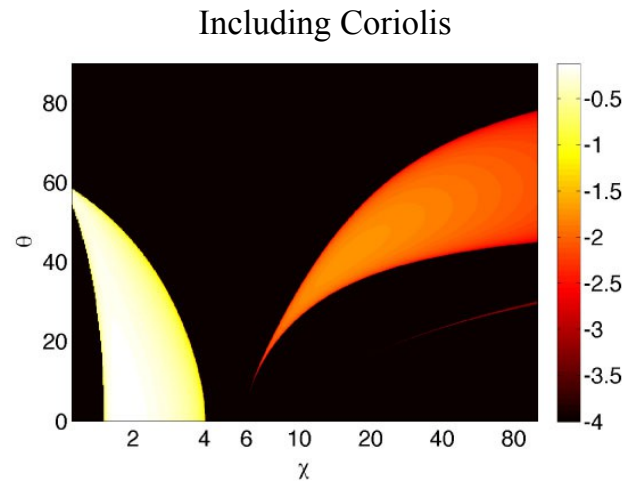
## Direct Cascade

Elliptic destruction occurs  
faster (turnover frequency)  
than viscous merging

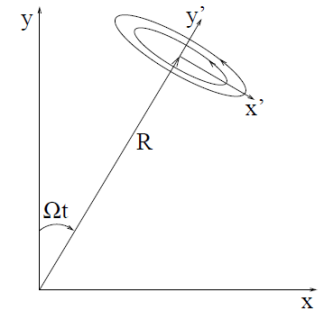
# Elliptic-Rotational Instability



Lesur & Papaloizou (2009)  
Bayly (1986)



Lesur & Papaloizou (2009)  
Miyakazi (1992)



## Instability of elliptic streamlines

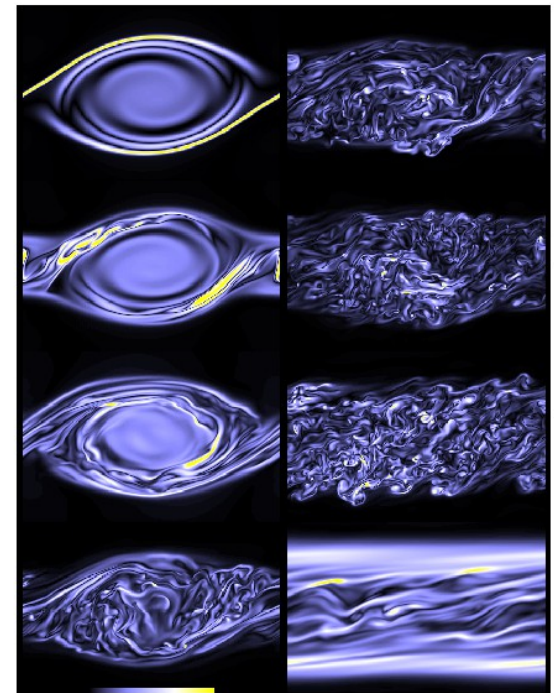
\* In the **non-rotating** case:

- **Resonance** between  
Strain field and Inertial waves

\* In the **rotating** case:

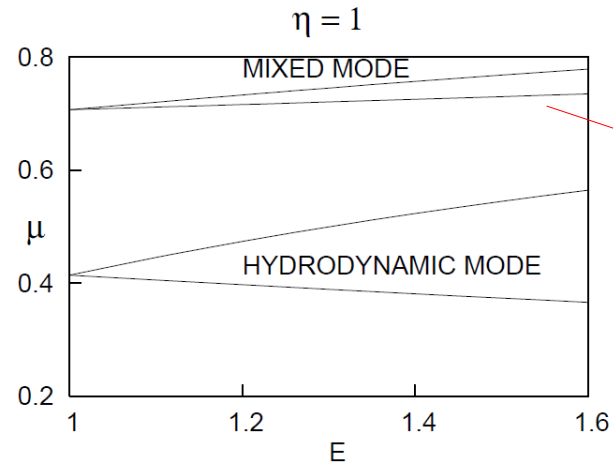
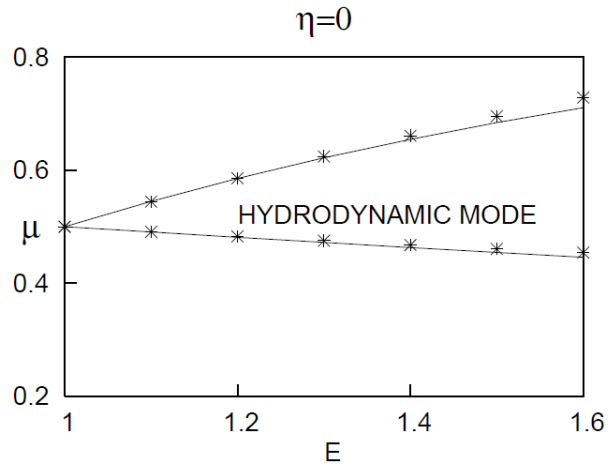
- Strong "horizontal" ( $\theta=0$ ) unstable mode:  
**Exponential growth of epicyclic disturbances**

Vortex coherence is destroyed  
Energy cascades forward and dissipates

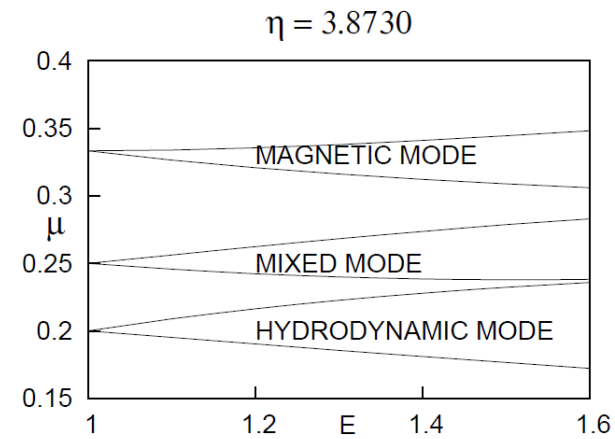
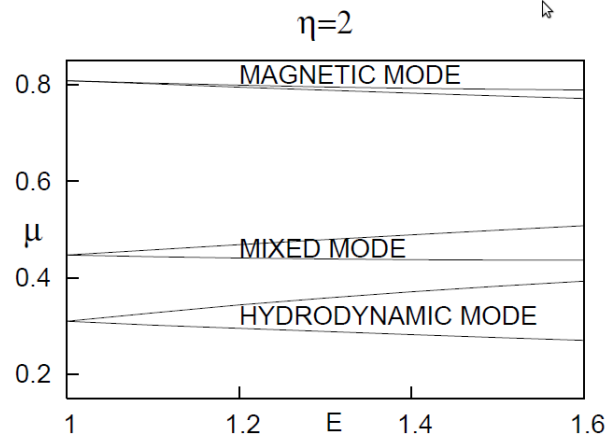


McWilliams (2010)

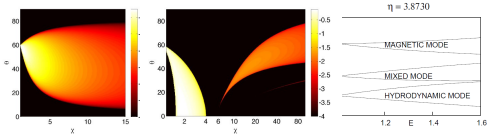
# Magneto-elliptic instability



*Kelvin + Alfvén wave*



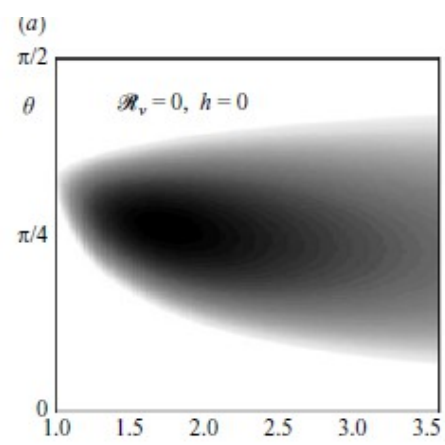
**B**



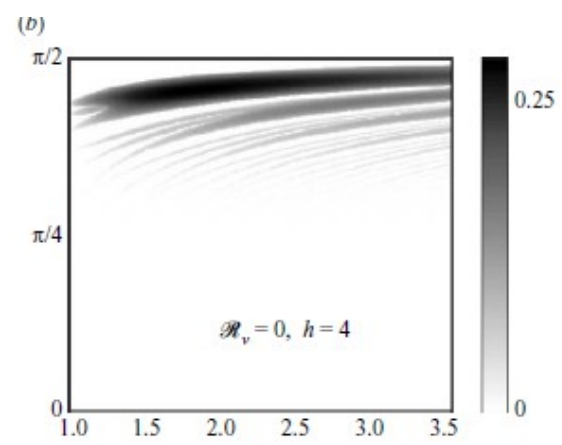
# Magneto-elliptic instability with rotation

$Rv$  = Inverse Rossby  
 $h$  = Magnetic field

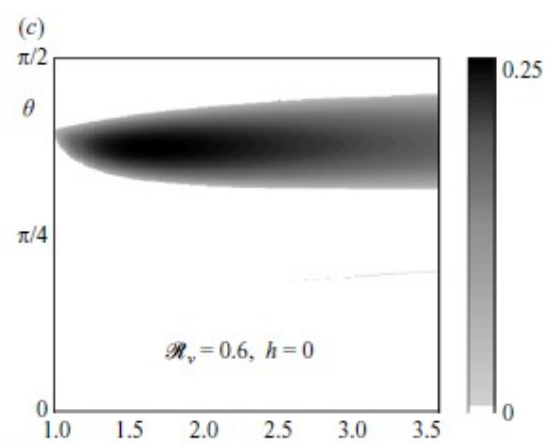
Pure elliptic



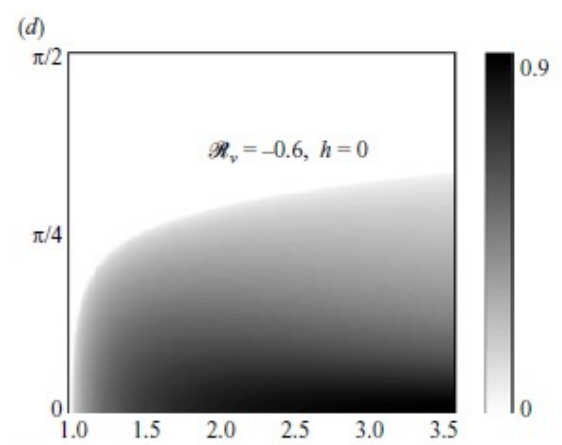
Pure magneto-elliptic



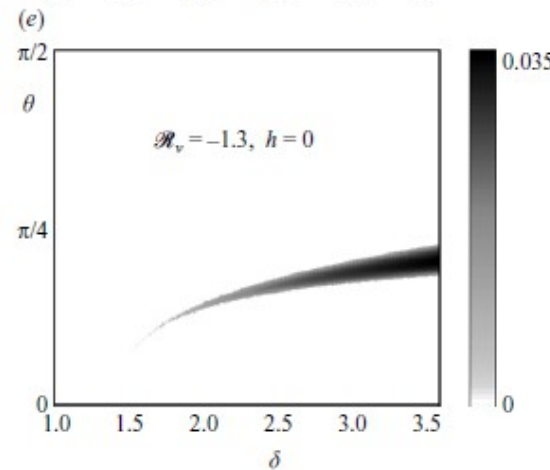
Elliptic-rotational  
cyclonic



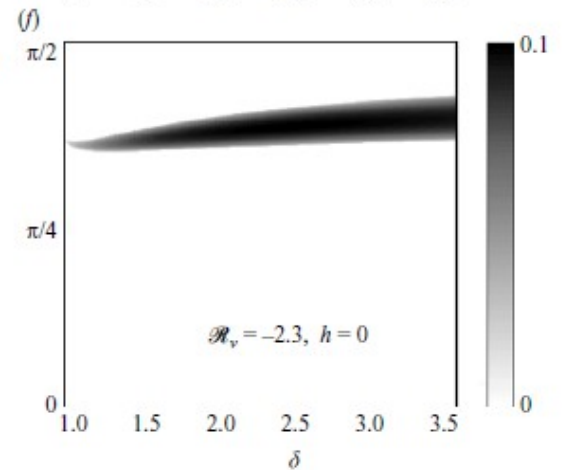
Elliptic-rotational  
anticyclonic



Elliptic-rotational  
Anticyclonic  
(slow rotation)

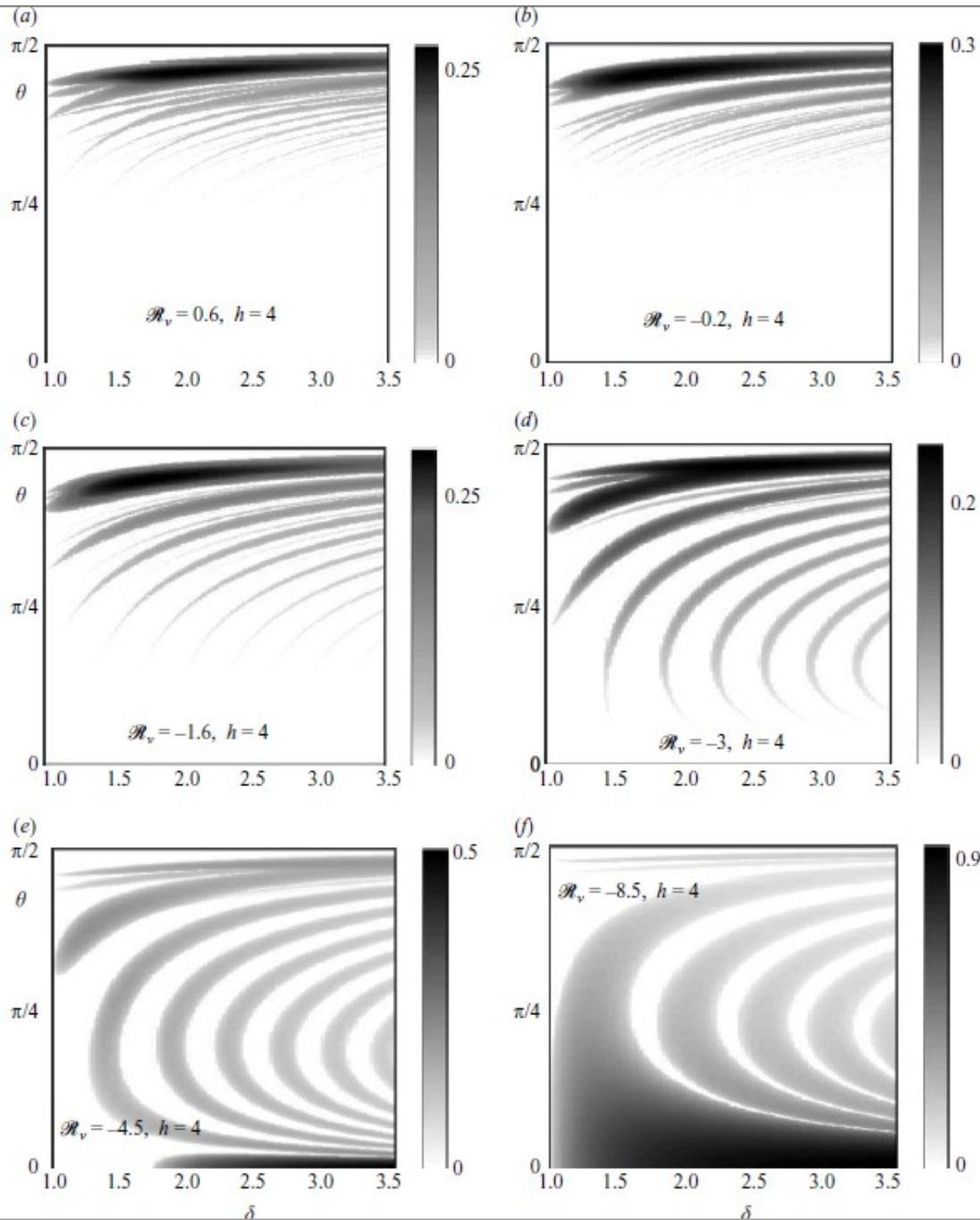


Elliptic-rotational  
Anticyclonic  
(slower rotation)



# Horizontal magneto-elliptic instability

$Ro$  = Inverse Rossby  
 $h$  = Magnetic field



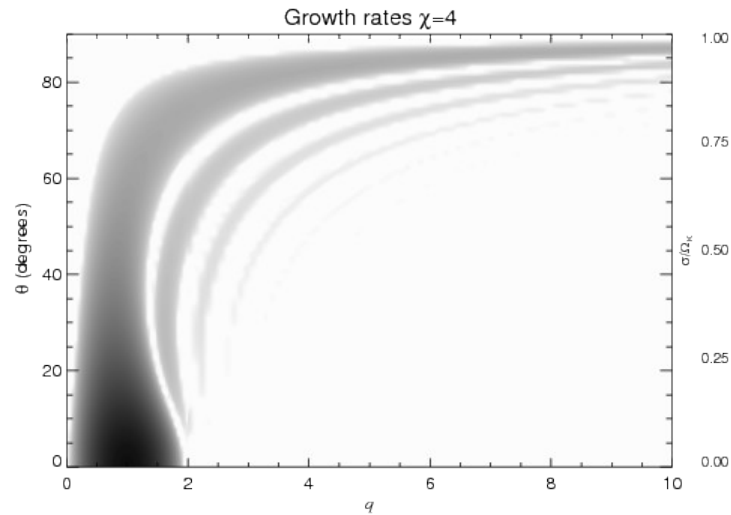
When rotation and field satisfy

$$Ro^{-1} < -\frac{h^2}{4}$$

a **strong horizontal mode** appears

Mizerski & Bajer (2009)

# Magneto-elliptic instability



Lyra & Klahr 2011  
(after Mizerski & Bajer 2009)

Mizerski & Bajer (2009, Journal of Fluid Mechanics)

“The presence of magnetic fields widens the range of existence of the horizontal instability to an unbounded interval of aspect ratios when

$$Ro^{-1} < -\frac{h^2}{4}$$

$$0 < k/k_{BH} < 2|Ro|^{1/2}$$

$$h = q / Ro$$

$$Ro = \frac{\Omega_V \delta}{\Omega_K}$$

$$\delta = \frac{1}{2}(\chi + \chi^{-1})$$

$$q = k / k_{BH}$$

$$k_{BH} = \frac{\Omega_K}{v_A}$$

# Magneto-elliptic instability

$$Ro = \frac{\Omega_V \delta}{\Omega_K}$$

$$\delta = \frac{1}{2}(\chi + \chi^{-1})$$

$\chi$  Aspect ratio

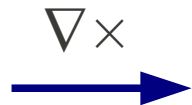
$$0 < k/k_{BH} < 2|Ro|^{1/2}$$

Write the criterion in terms of vorticity instead of angular frequency:

Vortex flow

$$u_x = -\Omega_V y / \chi$$

$$u_y = \Omega_V x \chi$$



Vorticity

$$\omega_T = \Omega_V (\chi + \chi^{-1}) = 2 \Omega_V \delta = 2 Ro \Omega_K$$

$$= \omega_V + \omega_{box}$$

$$= \omega_V - 3/2 \Omega_K$$

$$Ro = \frac{\omega_V}{2 \Omega_K} - \frac{3}{4}$$

In the no-vortex limit ( $\omega_V = 0$ ),  $Ro = -3/4$

$$0 < k/k_{BH} < \sqrt{3}$$



**Magneto-elliptic instability → No vortex limit**

$$0 < k / k_{BH} < \sqrt{3}$$

# Consistency

$$Ro = \frac{\Omega_V \delta}{\Omega_K}$$

$$\delta = \frac{1}{2}(\chi + \chi^{-1})$$

$\chi$  Aspect ratio

Kida solution

$$\Omega_V = -\frac{3\Omega_K}{2(\chi-1)}$$

$$Ro = -\frac{3}{4} \frac{\chi^2 - 1}{\chi(\chi - 1)}$$

$$\begin{aligned} \Omega_V &= 0 \\ \chi &\rightarrow \infty \end{aligned}$$

Remember:  
In the no-vortex limit ( $\omega_v=0$ ),  $Ro=-3/4$

$\chi \rightarrow \infty$

$$\lim_{\chi \rightarrow \infty} Ro = -\frac{3}{4}$$

Vortex flow

$$u_x = -\Omega_V y / \chi$$

$$u_y = \Omega_V x \chi$$

$\chi \rightarrow \infty$

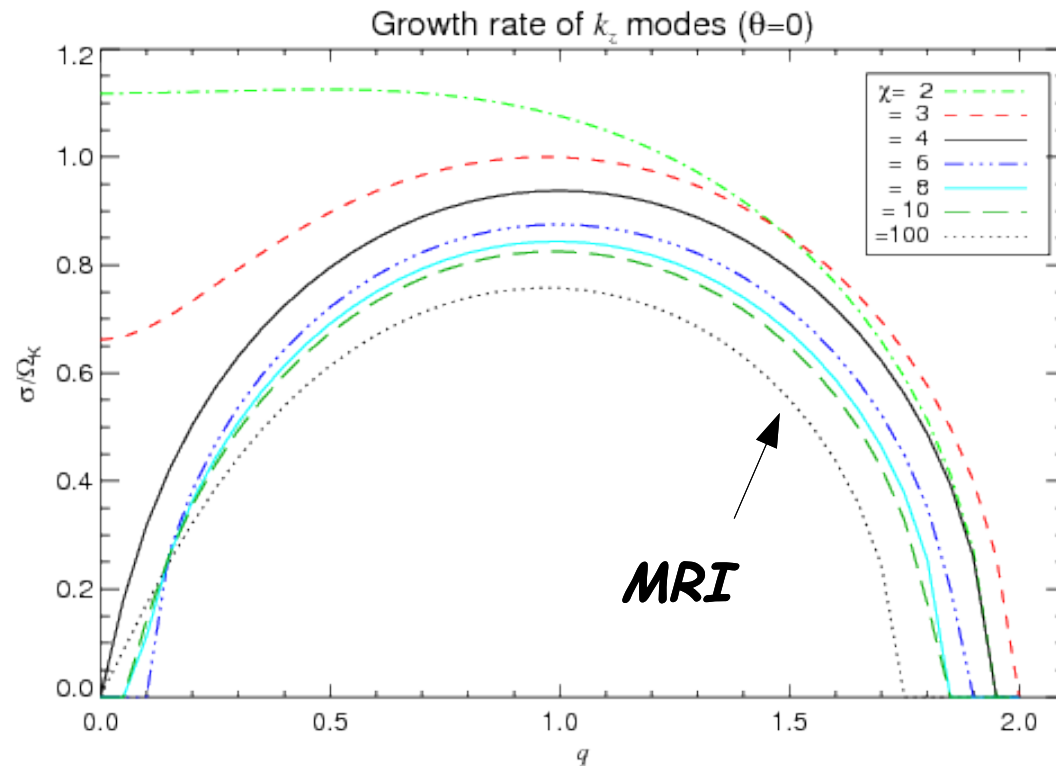
$$\lim_{\chi \rightarrow \infty} u_x = 0$$

$$\lim_{\chi \rightarrow \infty} u_y = -3/2 \Omega_K x$$

**QED!**

**A vortex of infinite aspect ratio is equivalent to a shear flow**

# Growth rates of the **Magneto-Elliptic-Rotational Instability**



... we explain both  
the MEI and the MRI  
as different manifestations  
of the same  
**Magneto-Elliptic-Rotational  
Instability**

Mizerski & Lyra (2012)

## On the connection between the magneto-elliptic and magneto-rotational instabilities

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<sup>3</sup>Jet Propulsion Laboratory, 4800 Oak Grove Drive, Pasadena CA 91109, USA

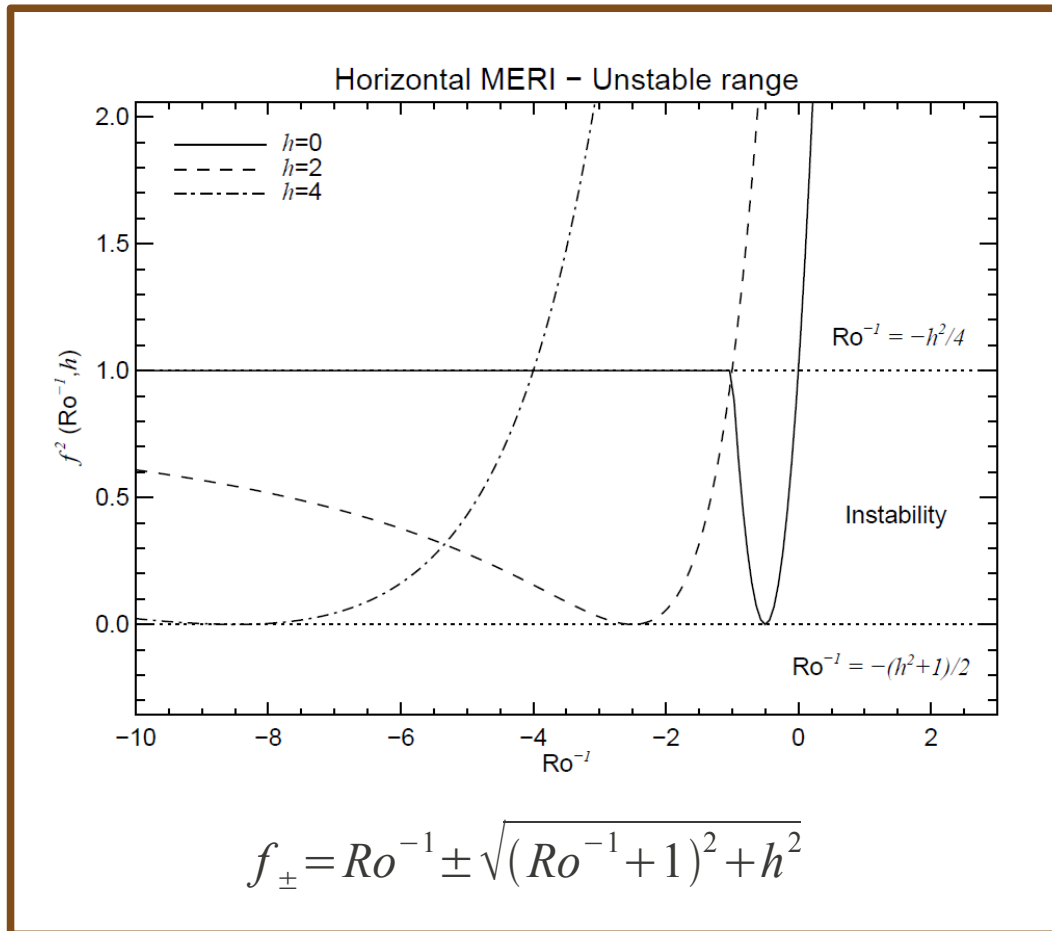
<sup>4</sup>NASA Carl Sagan Fellow

(Received 11 July 2011; Accepted 15 February 2012.)

# Dispersion Relation

$$\begin{bmatrix} \hat{v}_x \\ \hat{v}_y \\ \hat{b}_x \\ \hat{b}_y \end{bmatrix} = \begin{bmatrix} 0 & 1 + \varepsilon + 2\text{Ro}^{-1} & ih & 0 \\ -(1 - \varepsilon) - 2\text{Ro}^{-1} & 0 & 0 & ih \\ ih & 0 & 0 & -(1 + \varepsilon) \\ 0 & ih & 1 - \varepsilon & 0 \end{bmatrix} \begin{bmatrix} \hat{v}_x \\ \hat{v}_y \\ \hat{b}_x \\ \hat{b}_y \end{bmatrix}$$

Mizerski & Lyra (2012)



## Horizontal MEI

$$\frac{\sigma^2}{\gamma^2} = \varepsilon^2 - \text{Ro}^{-2} \left[ \sqrt{(\text{Ro}^{-1} + 1) + q^2} - 1 \right]^2$$

$$\varepsilon = 1 \quad \downarrow \quad \text{Ro} = -3/4$$

## Horiz. MEI at shear flow lim.

$$\frac{\sigma^2}{\gamma^2} = \frac{8}{9} \left( \sqrt{16q^2 + 1} - 2q^2 - 1 \right)$$

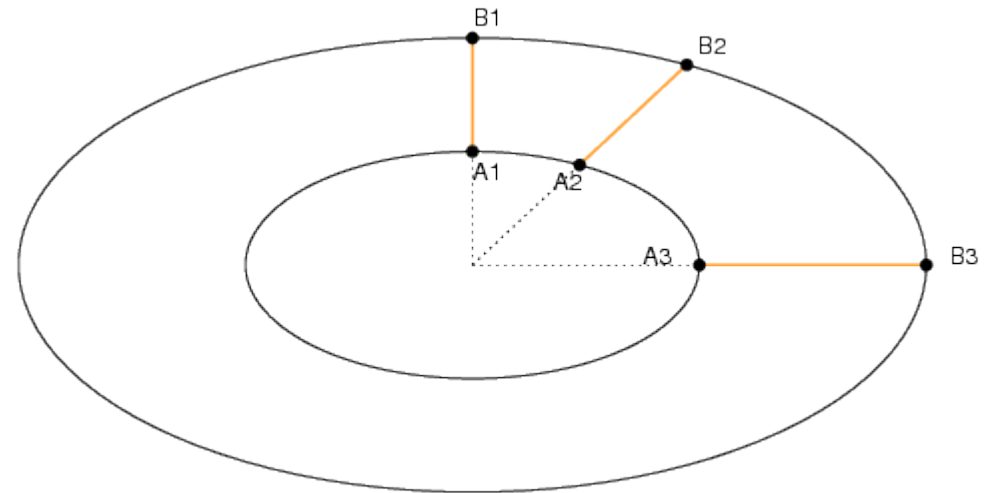
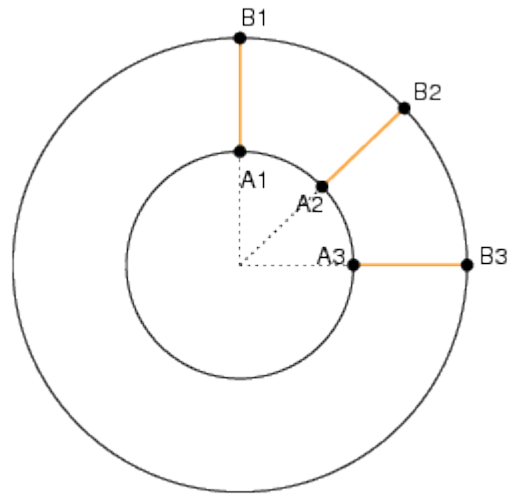
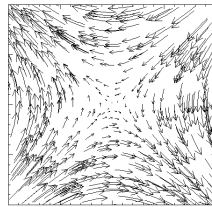
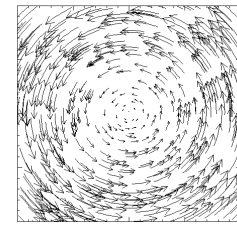
### Properties

$$0 < q < \sqrt{3}$$

$$\sigma = \gamma = 3/4 \Omega$$

$$q = \sqrt{15}/4 \approx 0.9682$$

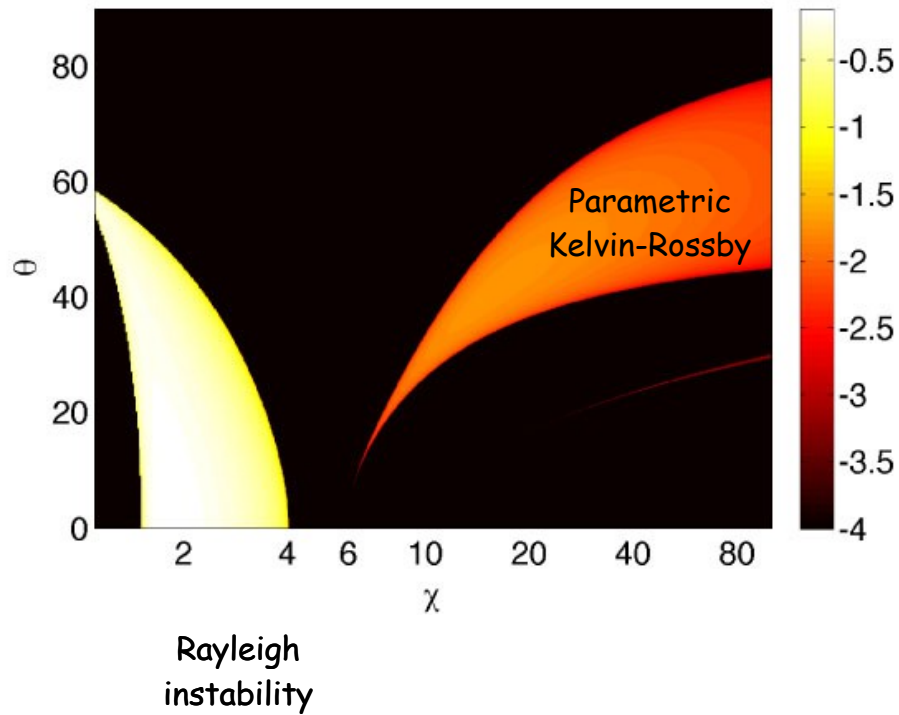
# Common ground between MRI and MEI



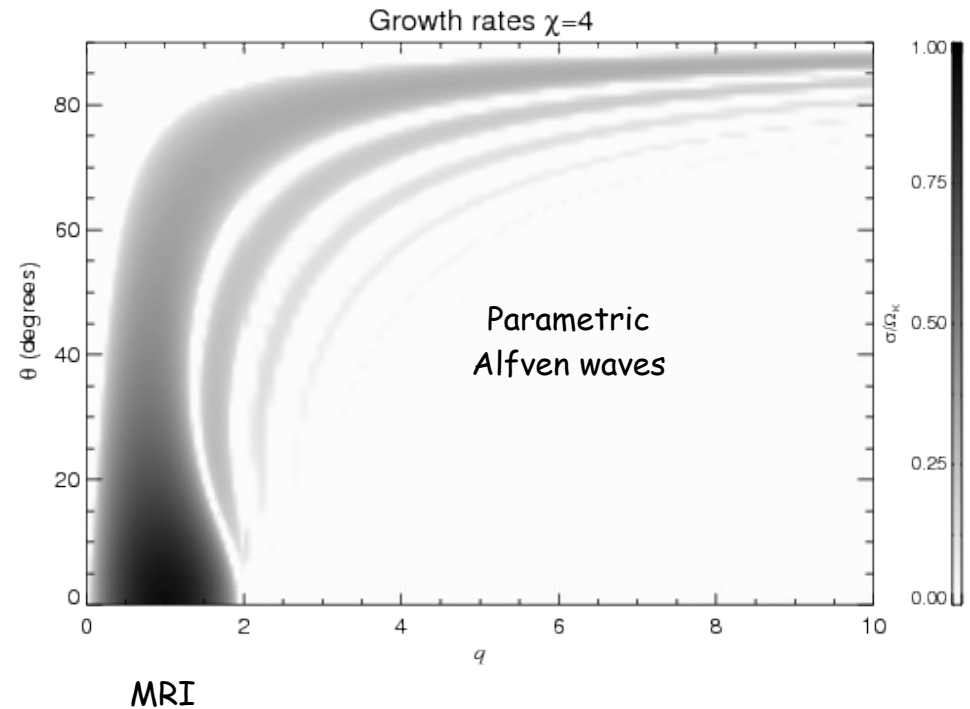
Elliptic streamlines have shear  
even in uniform rotation.

# Destroying vortices with 3D instabilities

## Elliptic-Rotational



## Magneto-Elliptic-Rotational

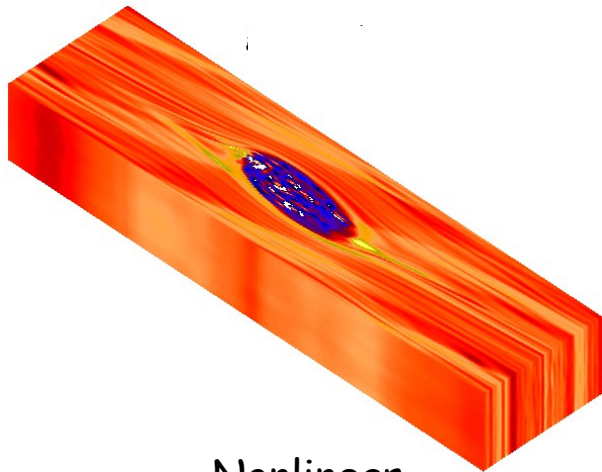


# Sustaining vortices

Mechanisms to  
*inject vorticity*

to counteract the vorticity lost in the direct cascade

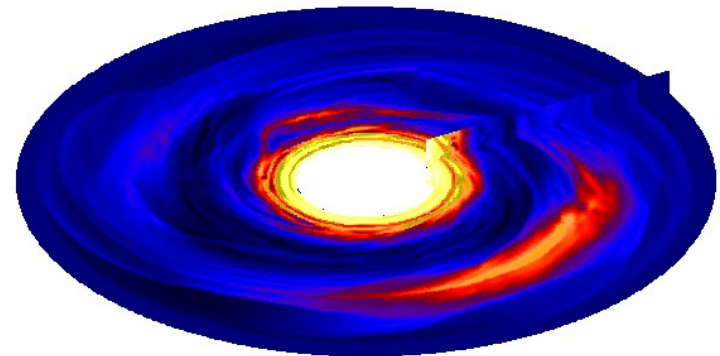
## Baroclinic Instability



Nonlinear

Powered by:  
buoyancy, thermal diffusion  
(baroclinic source term)

## Rossby Wave Instability



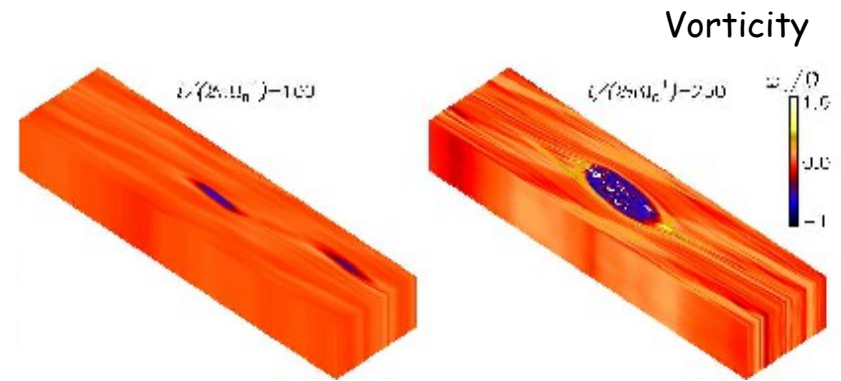
Linear

Powered by:  
Modification of shear profile  
(*external vortensity reservoir*)

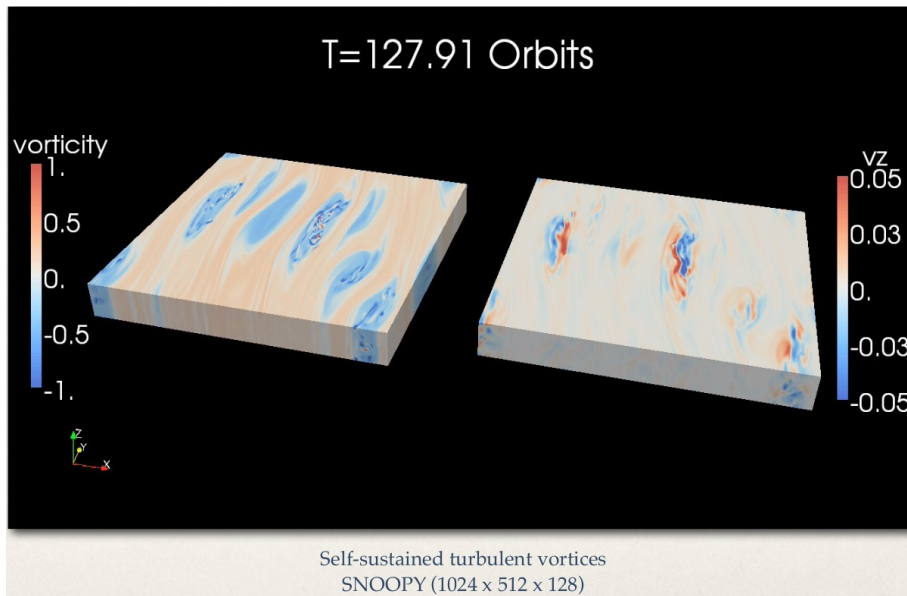
# Baroclinic Instability and Elliptic Instability

Despite the elliptical instability, **baroclinity keeps the vortex coherent.**

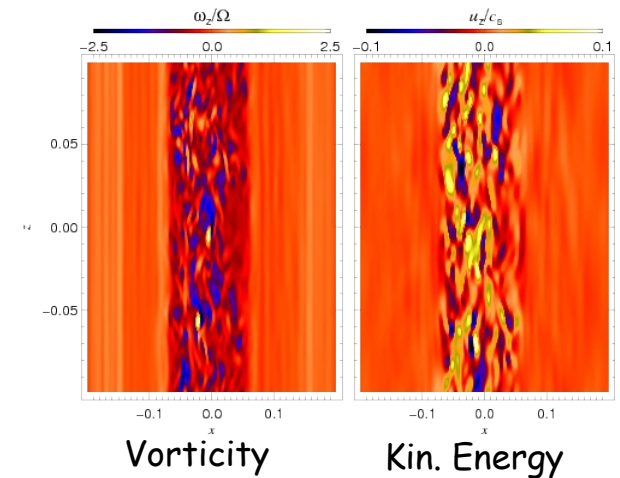
The result is "core turbulence" only



Lyra & Klahr (2011)



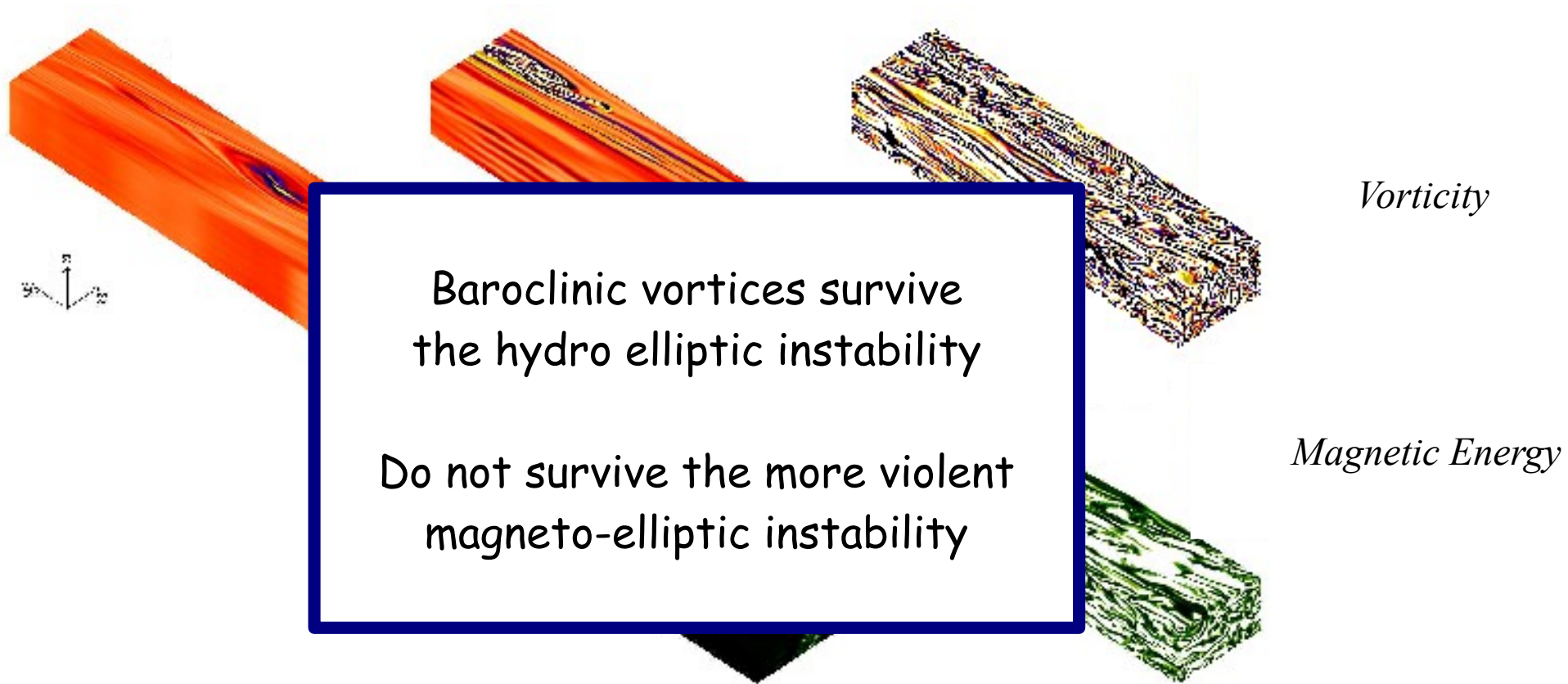
Lesur & Papaloizou (2010)





# Baroclinic Instability and Magneto-Elliptic Instability

What happens when the vortex is magnetized?

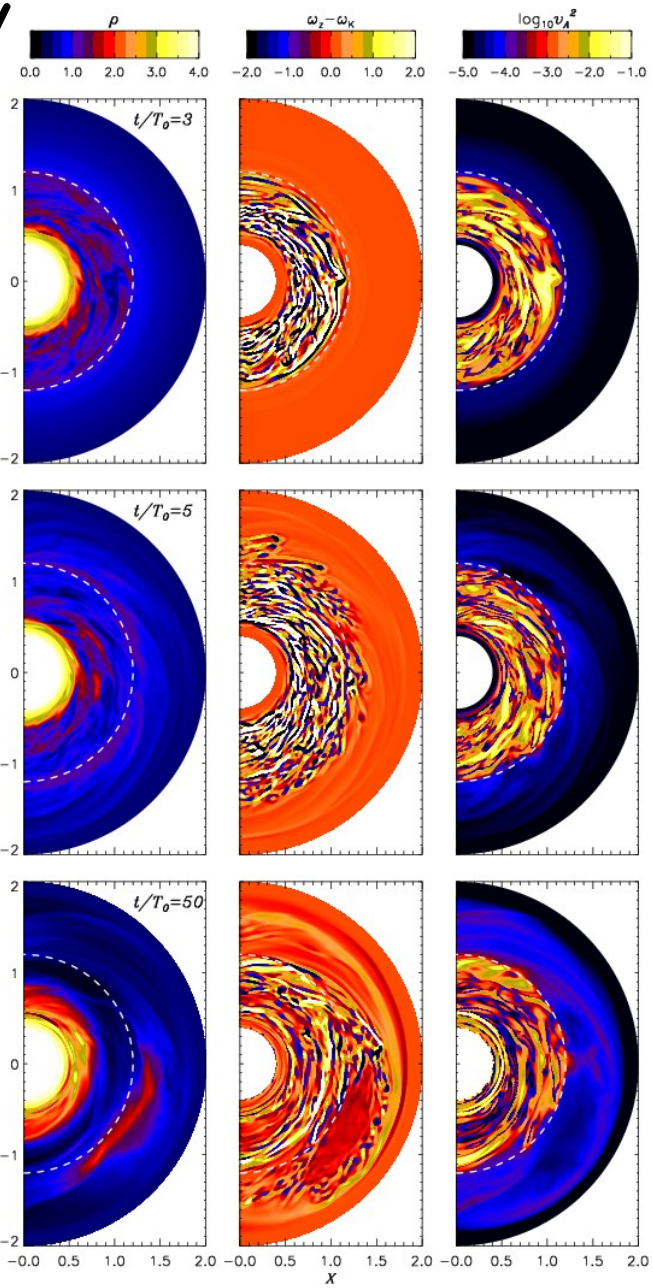
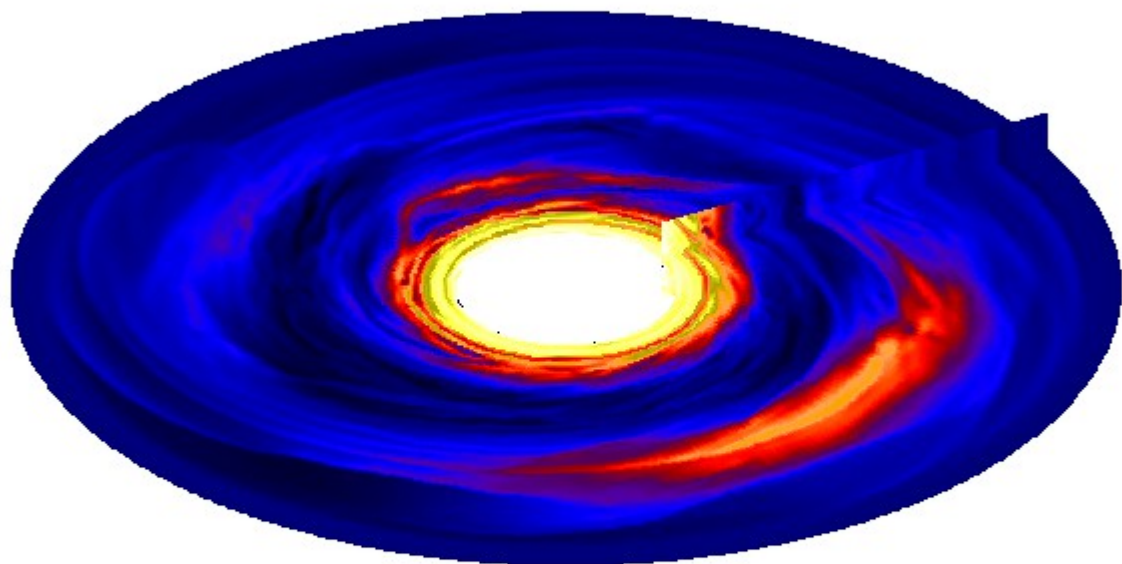


Vortex gone!

Lyra & Klahr (2011)

# Rossby Wave Instability

$t = 22.28 \tau_0$



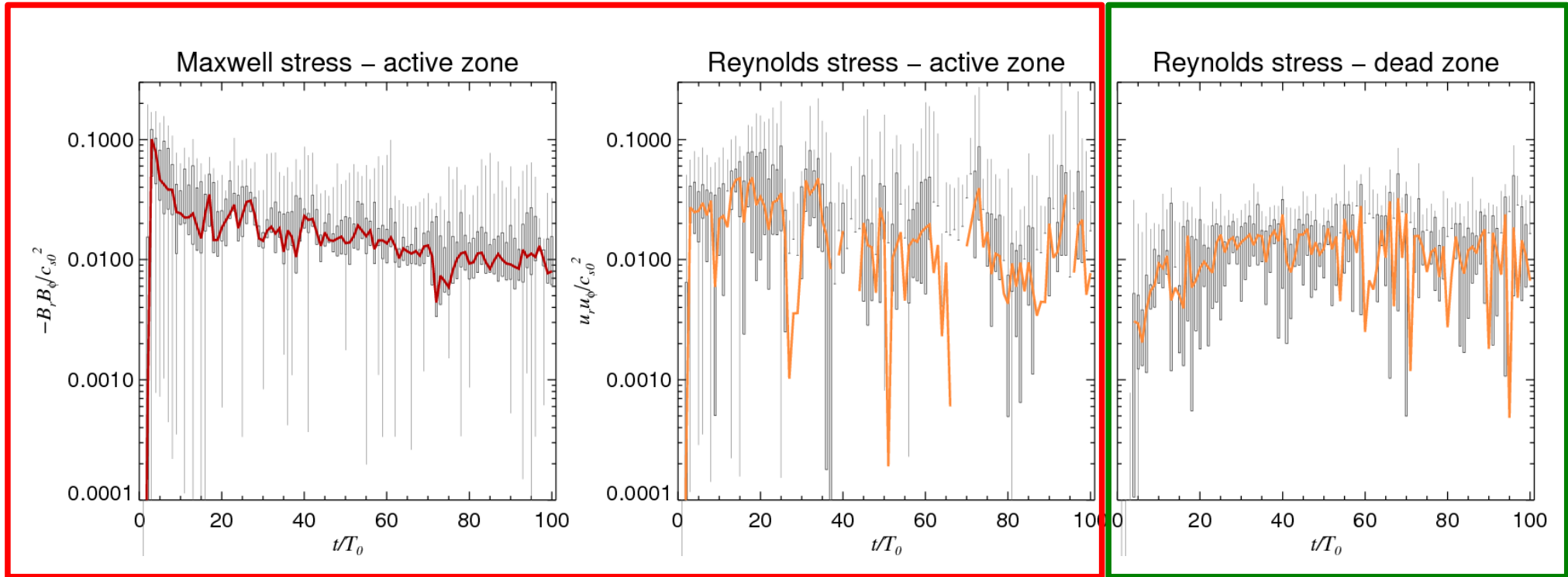
Magnetized inner disk + resistive outer disk

Lyra & Mac Low (2012)

# Significant angular momentum transport

Active zone

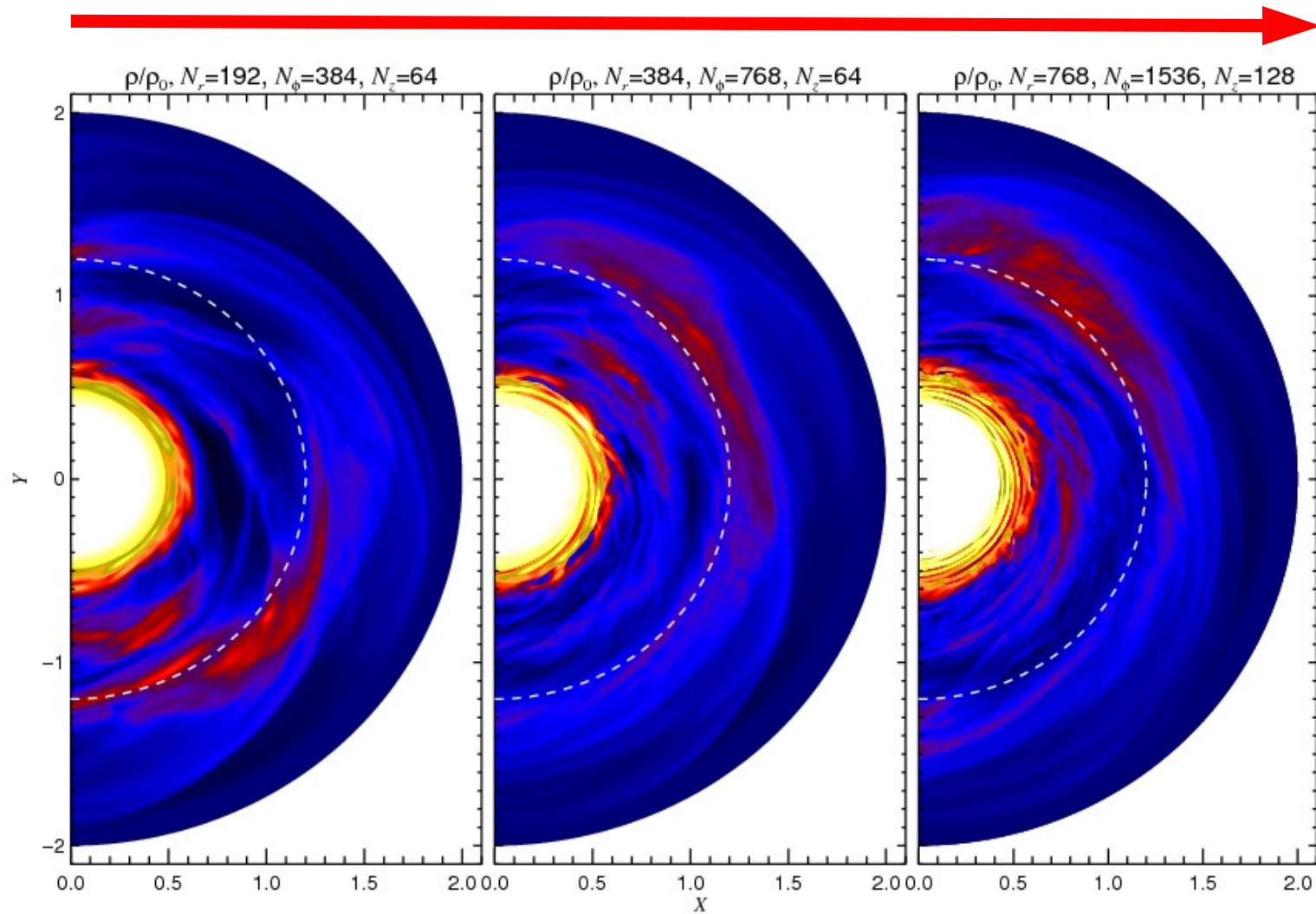
Dead zone



Large mass accretion rates in the **dead zone**,  
comparable to the MRI in the **active zone**!

# Fishy vortex in the active zone...

Resolution

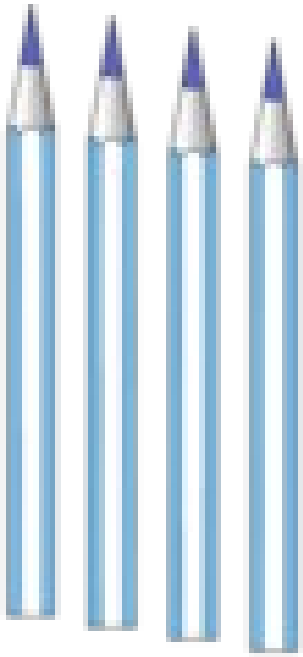


$\Delta r/H = 10$

$\Delta r/H = 20$

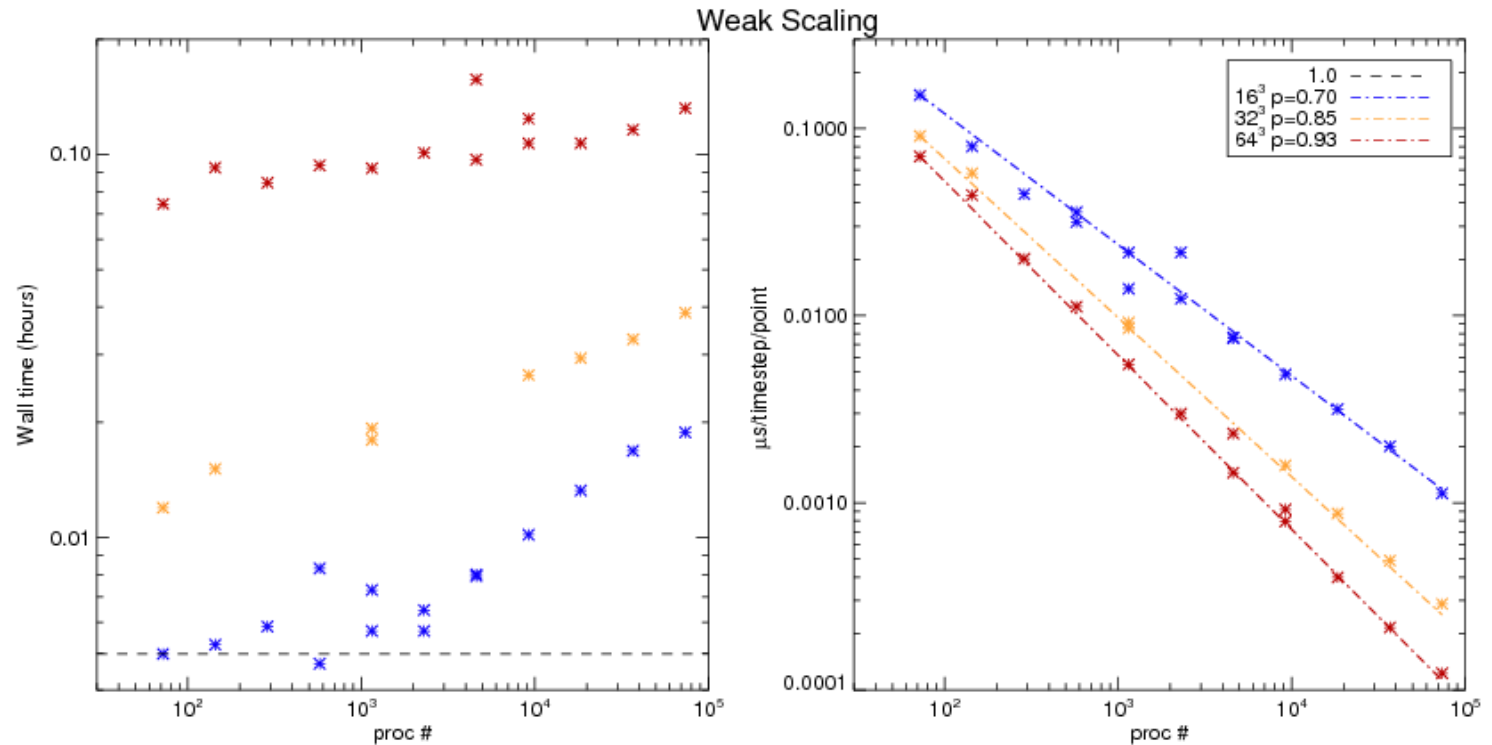
$\Delta r/H = 40$

# High end computing



## The Pencil Code

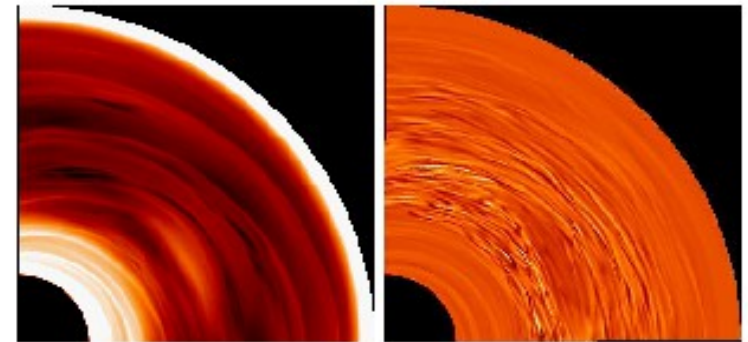
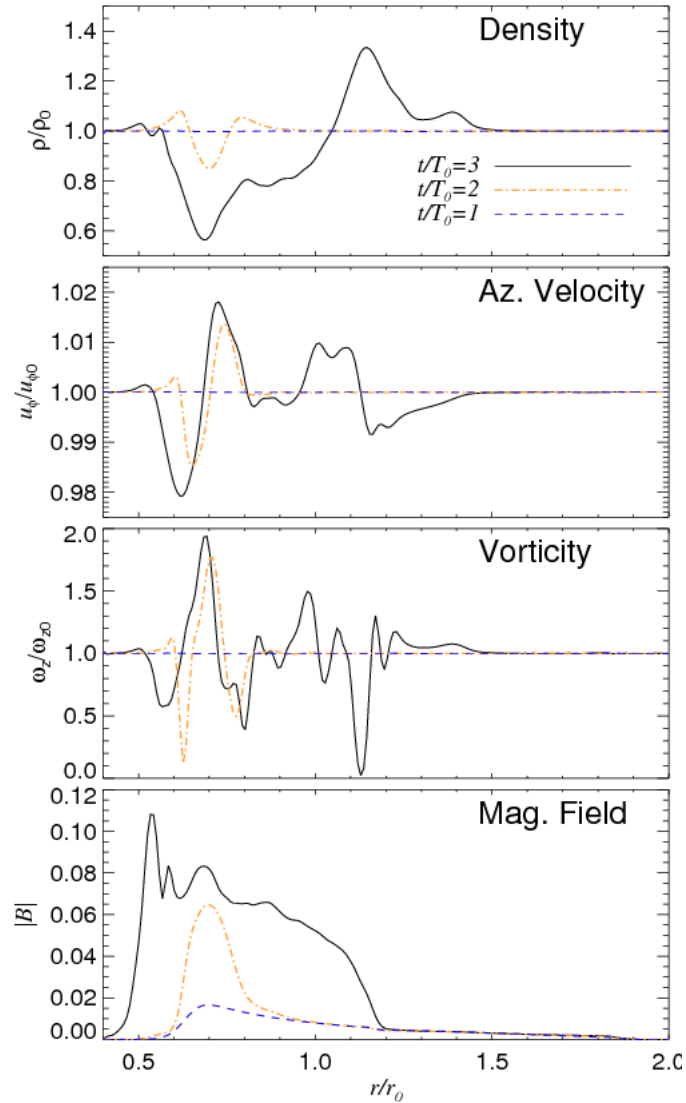
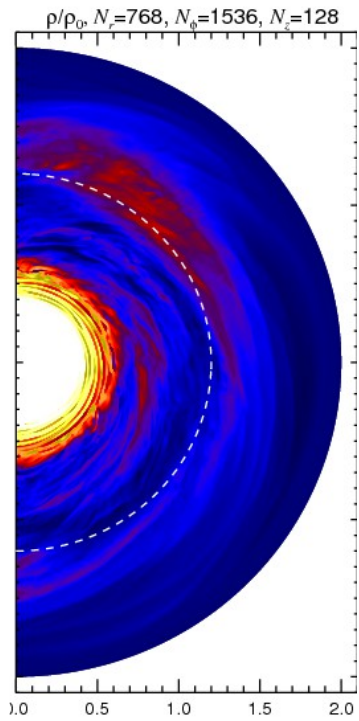
Brandenburg & Dobler (2002)



Good scaling up to > **70,000 processors !**  
(At NICS - Kraken)

# A zonal flow?

(see also R. Lovelace's talk)



density

vorticity

Fromang & Nelson (2005)

Lyra & Mac Low (2012)

## Facts

Rossby vortices survive  
the magneto-elliptic instability,  
whereas baroclinic vortices do not.

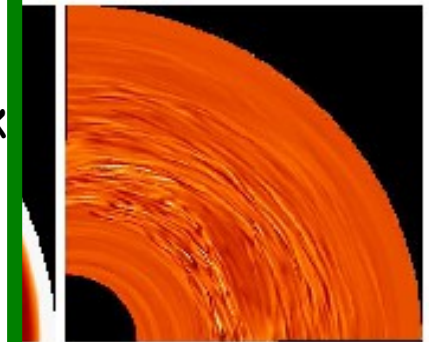
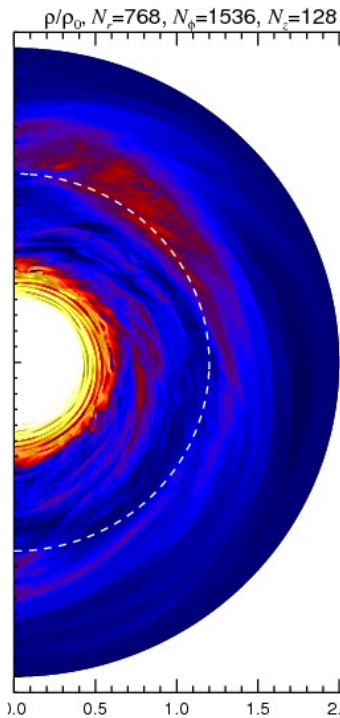
"Vortex survival is a balance  
between production and destruction"  
- John Papaloizou's talk

## Conjecture

Active zone Rossby vortex survives because RWI  
produces vorticity faster than the MEI destroys,  
whereas BI does not.

Strength of vorticity injection/destruction

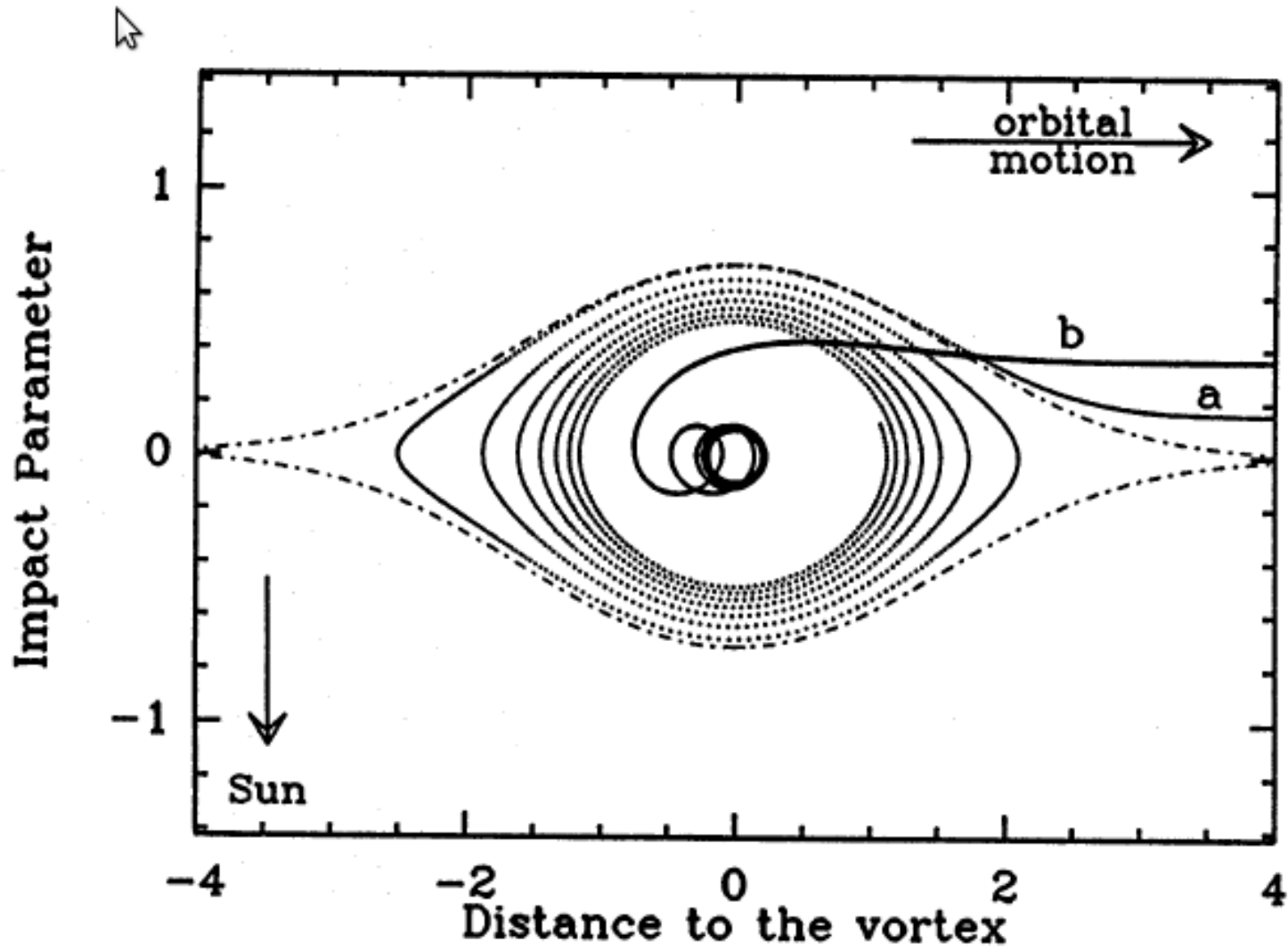
$$EI < BI < MEI < RWI$$



vorticity

Nelson (2005)

# Forming Planets



Barge & Sommeria (1995)

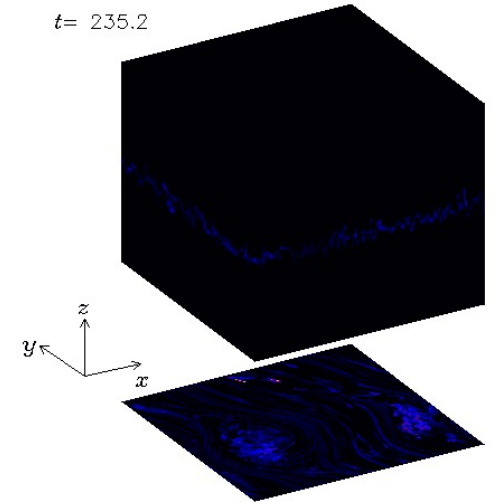
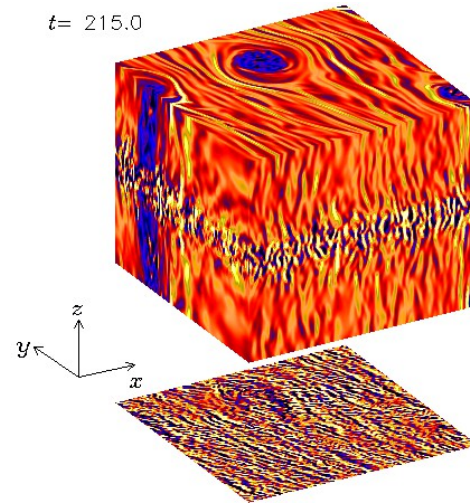


# Forming Planets

Gas

Particles

Backreaction



Vortices are  
**not destroyed**  
by heavy particle load.

No  
Backreaction

